

## Empirically adjusted ray amplitudes in the vicinity of the critical region

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### Resumen

Se desarrolla un método para corregir empíricamente las amplitudes de las ondas P producidas a partir de la teoría geométrica de rayos en la vecindad de la distancia crítica  $x_c$ , donde es bien sabido que la teoría produce amplitudes erróneas. Se encuentra que simplemente mediante un suavizamiento de las amplitudes dentro de una zona definida como "zona crítica" con un polinomio de tercer grado, las curvas de AVO generadas mediante la teoría de rayos se ajustan muy bien a los resultados exactos. El método toma en consideración implícitamente las variaciones de amplitud vs. offset, frecuencia, altura de la fuente y receptor por encima de las interfaces y de los parámetros físicos del modelo geológico.

### Abstract

An empirical approach for correcting geometrical ray theory P-wave amplitudes in the vicinity of the critical distance  $x_c$ , is presented. Ray theory is well known to produce inaccurate amplitudes. It is found that simply smoothing the amplitudes within the previously defined "critical zone" with a third-degree polynomial provides an excellent match of the AVO curves generated by the ray theory with the exact results. The method implicitly takes into account the amplitude variation vs. offset, frequency, source-receiver height above the interface and physical parameters of the geological model.

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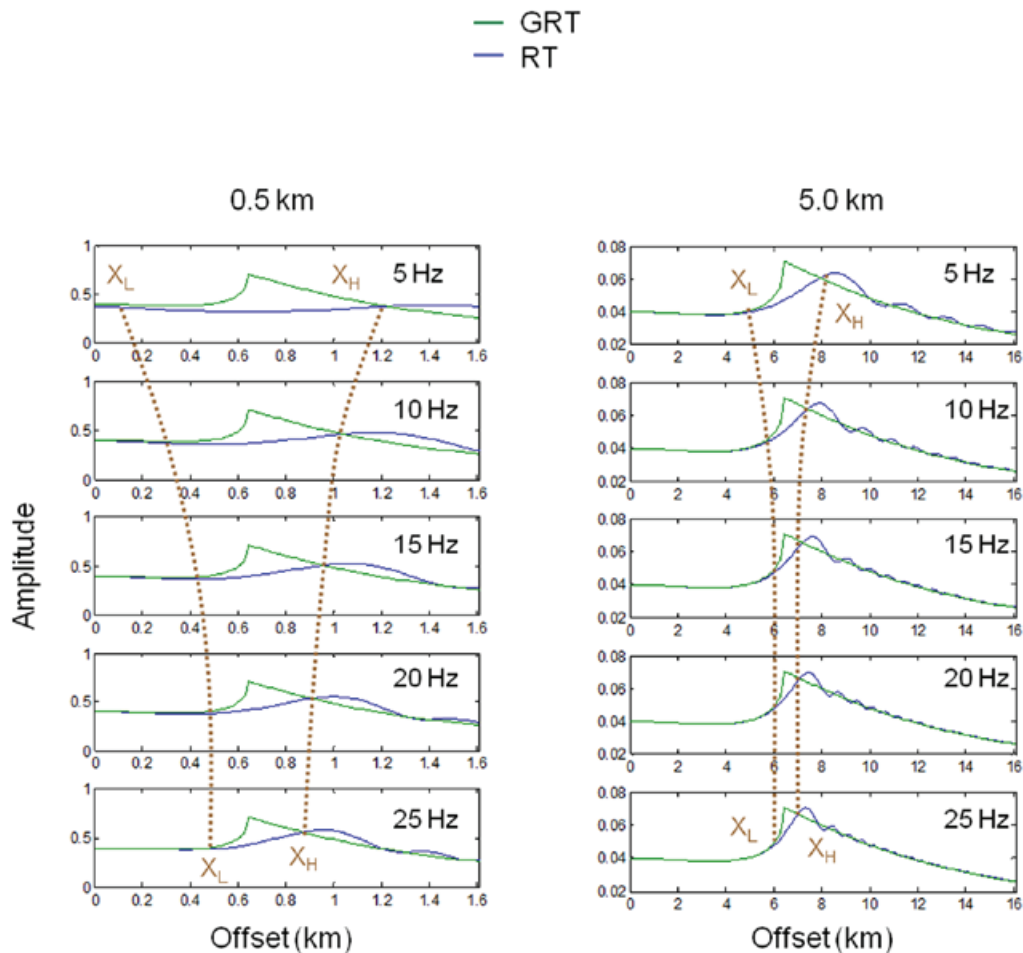
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## Introduction

Forward seismic modeling is used in multiple applications to predict the amplitude behavior of physical models, for oil exploration, earthquake seismology and other applications. It is typically based on grid methods that discretize the wave equation and the physical model; another approach is deriving high-frequency approximations of the wavefield based on geometrical optics as provided by ray tracing methods. The latter method delivers quick acceptable results for a variety of geometrically complex geological models; the band-limited wavefield is simulated by seismic rays. The grid numerical solutions are often computationally demanding as the equations must be solved in each of the numerical cells.

The Ray tracing methods, provide a good approximation for models in which the size of the heterogeneities is larger than the dominant wavelength: it is a high frequency solution. The solution, in terms of amplitudes and phases, is remarkably good for offsets smaller than the critical distance (pre-critical offsets) and moderately so in the post-critical zone. The amplitudes however completely diverge around the critical distance  $x_c$  due to singularities at caustics and foci (Cerveny, 1966a, 1966b). This window of divergent amplitudes can be called "critical region". Several methods have been developed to correct ray amplitudes in this region. In flat-layered media the standard approach is to use Weber-Hermite functions (Cerveny and Ravindra, 1971). Another approach is to use Gaussian-beam summation method (Norris, 1986).



**Figure 1.** AVO curves for 5 different frequencies comparing generalized ray (GRT) and ray theory (RT). Brown dotted lines show how the width of the critical zone  $x_z$  varies exponentially with frequency and source receiver height above interface (1 and 5 km). The physical parameters of the model can be found in Table 1.

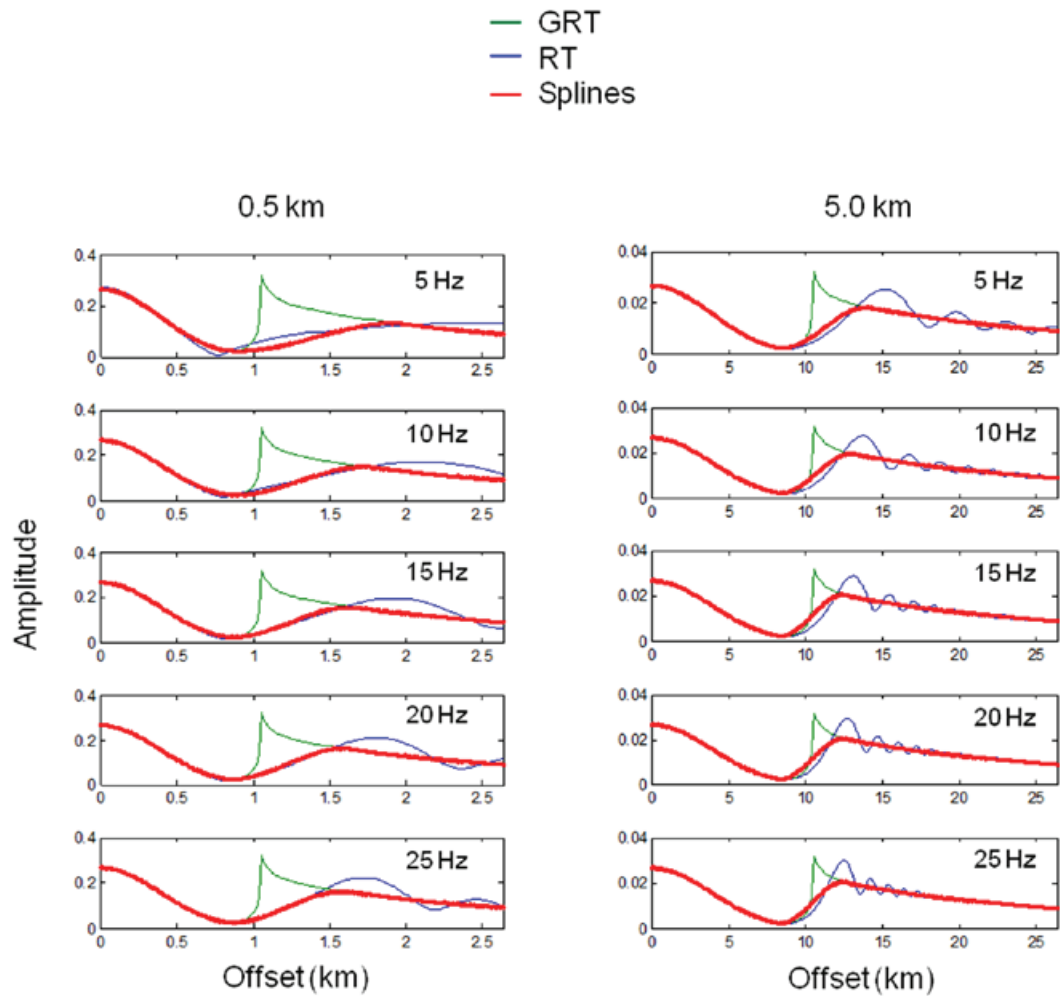
**Table 1.** Physical parameters of the geological model (see Figure 1).

	$\rho(\text{kg/m}^3)$	$V_p(\text{km/s})$	$V_s(\text{km/s})$
Upper half-space	2.0	2.0	1.3
Lower half-space	2.5	3.7	1.5

The objective of this work is to develop a fully empirical approach, which is completely independent of any seismic theory describing the propagation of waves. To achieve this, it is shown in the following section that simply by defining a 'critical region' and smoothing the AVO curves within it, the amplitude is properly corrected. Here the problem is restricted to two elastic isotropic half-spaces in welded contact for incident and reflected P waves, but it can easily be generalized to n-layers. The results are compared to the exact solution based on the generalized ray theory (Kanasewich *et al.*, 1983).

### Method

To quantify the behavior of amplitude errors within the critical region, comparisons are made in this work between the results obtained with an exact analytical method, i.e., Generalized Ray Theory (GRT), (Kanasewich *et al.*, 1983) and Geometrical Ray Theory (RT) for realistic physical parameters in each of the two half-spaces. Of all trials, we include only the most representative cases: Poisson's ratio in each of the two half-spaces varying, from 0.1 to 0.4, varying Poisson's ratio contrast between both



**Figure 2.** AVO curves comparing the exact analytical amplitudes (GRT-blue), geometrical ray theory amplitudes (RT-green) and corrected ray theory amplitudes (CRT-ray) for different frequencies (5-15 Hz) and two different source-receiver height above the interface (0.5 and 5 km). The correction involves the amplitude smoothing within the critical zone  $xz$ . The physical parameters of the model can be found in Table 2.

**Table 2.** Physical parameters of the geological model (see Figure 2).

	$\rho(\text{kg/m}^3)$	$V_p(\text{km/s})$	$V_s(\text{km/s})$
Upper half-space	2.0	2.9	1.3
Lower half-space	2.5	4.0	2.6

half-spaces, from 0.36 to 3.0, increasing velocity contrast  $Vp_1/Vp_2$  from 0.5 to 0.967, varying source-receiver height above the interface, from 0.5 to 5 km, and varying frequency, from 5 to 25 Hz. The GRT method makes use of a Cagniard-Pekeris method by decomposing the wavefield into generalized rays.

From the seismograms in Fig. 1, it is consistently observed that the main difference between the two results is in terms of the amplitudes in the AVO curves within the critical zone. The phase differences are negligible for this elastic isotropic case. It is expected however that anelasticity could yield significant phase differences as well.

Our approach comprises two steps. The first step consists of defining the critical zone  $xz$  where the amplitudes are to be interpolated. This is done by finding the lower and higher bound  $x_L$  and  $x_H$ ; they correspond to the smallest and largest offsets about  $x_c$  where the RT amplitudes start to diverge from the exact results (GRT). Figure 1 shows that these boundaries vary exponentially with (1) the physical parameters of the model, (2) the frequency and (3) the source-receiver height above the interface.

It should be noted that the differences in amplitudes between RT and GRT within  $xz$  diminish with increasing frequency and with increasing source-receiver height above the interface.

The most accurate way to determine these bound is to compare the exact (GRT) and the RT AVO curves only at the lowest and highest frequencies of the wavefield to get two sets of values of  $x_L(f_0, \infty)$  and two of  $x_H(f_0, \infty)$  and then fit an exponential function to get the remaining  $x_L(f)$ 's and  $x_H(f)$ 's across the entire frequency range. This is only done once for each geological model;  $f_0, \infty$  stands for zero and infinite frequencies. Since the exponential function of frequency of  $x_L$  and  $x_H$  corresponds to a linear function of frequency in the log scale, the only difference among geological models is the slope and intercept of  $x_L(f)$ 's and  $x_H(f)$ 's.

To first order it turns out that  $x_L$  is smaller than the critical distance  $x_c$  whose first order derivative of the AVO curve undergoes the first polarity reversal. As these first-order derivatives need not be accurate, they are computed numerically with a first-order accurate central difference algorithm  $y'(x) = [y(x+h) - y(x-h)] / h$ . Here  $y$  represents the wave amplitude at offset  $x$  (receiver location),  $y'$  to its corresponding first order derivative and  $h$  to the receiver interval. These derivatives are computed only for offsets smaller than the critical distance.

As for the upper bound of the critical zone  $xH$ , it is observed, again to first order, that  $x_c < xH' < 2x_c$  where  $xH = xH'x_c$ . In practice,

$$x_{H'} = [Vp_1/Vp_2 + 0.35x_c] / h \quad (1)$$

independently of the source-receiver height above the interface.

The second step of the empirical approach to correct RT AVO within  $xz$  is to determine the type of interpolation to be used. It is found that a cubic polynomial is best suited for this purpose as the rate of change of amplitudes varies rapidly near  $x_L$  and  $x_H$ , especially the latter. Cubic splines

$$y = Ay_j + By_{j+1} + Cy''_j + Dy''_{j+1} \quad (2)$$

have a smooth first derivative and a continuous second derivative both within the interval and at its boundaries, which makes them amenable to smooth the ray theory amplitudes. The  $y''$ 's correspond to the second derivatives of the interpolation polynomial  $y$ , and the coefficients are given by

$$A \equiv \frac{x_{j+1} - x}{x_{j+1} - x_j} \quad B \equiv 1 - A = \frac{x - x_j}{x_{j+1} - x_j}$$

$$C \equiv \frac{1}{6}(A^3 - A)(x_{j+1} - x_j)^2 \quad D \equiv \frac{1}{6}(B^3 - B)(x_{j+1} - x_j)^2$$

Using Eq (2) it can be shown that the system of equations in the interval  $j = 2, \dots, N-1$  ( $xz$ ) is just

$$\frac{x_j - x_{j-1}}{6} y''_{j-1} + \frac{x_j - x_{j-1}}{6} y''_j + \frac{x_{j+1} - x_j}{6} y''_{j+1} = \frac{y_{j+1} - y_j}{x_{j+1} - x_j} - \frac{y_j - y_{j-1}}{x_j - x_{j-1}} \quad (3)$$

This system leads to  $N-2$  linear equations in  $N$  unknowns,  $y''_i = 1, \dots, N$  which is thus completed with boundary conditions at  $x_1$  and  $x_N$  for a unique solution. This is accomplished by setting the derivatives at the outer bounds to  $y''_1 = y''_N = 0$ , yielding the so-called *natural* cubic splines. The spatial locations  $x_1$  and  $x_N$  correspond to  $x_L - h$  and  $x_H + h$ . Thus, the algorithm first computes the second derivatives of the RT AVO curve and solves the tridiagonal system to obtain the cubic-spline interpolated amplitudes within the critical zone.

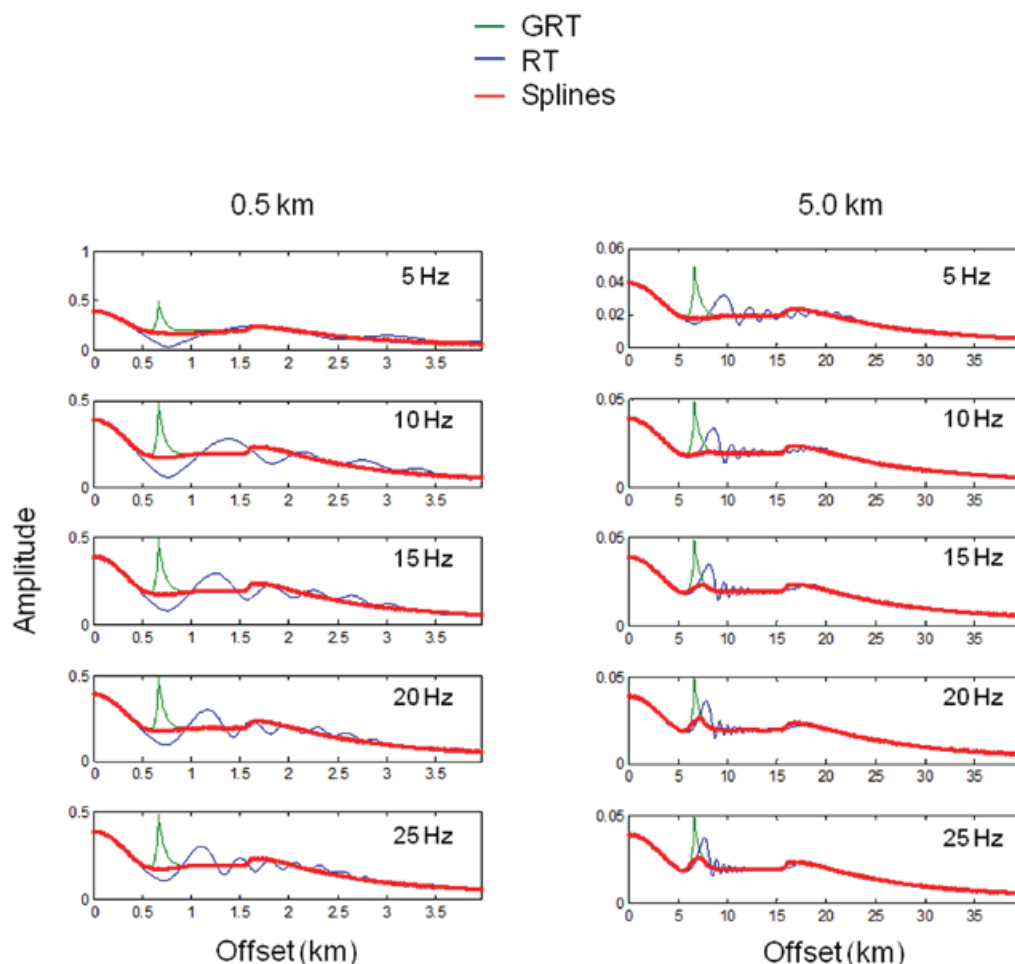
## Numerical results

Figures 2 through 4 show the corrections in RT amplitudes within the critical zone using the empirical approach proposed here. The physical parameters from which the AVO curves are computed can be found in Tables 2-4. As can be observed in these figures, the fit to the exact analytical amplitudes is very good. The most critical part of the procedure is to determine accurately the window size of the critical zone.

## Conclusions

We present a fully empirical approach to correct the ray theory P-wave amplitudes within the critical

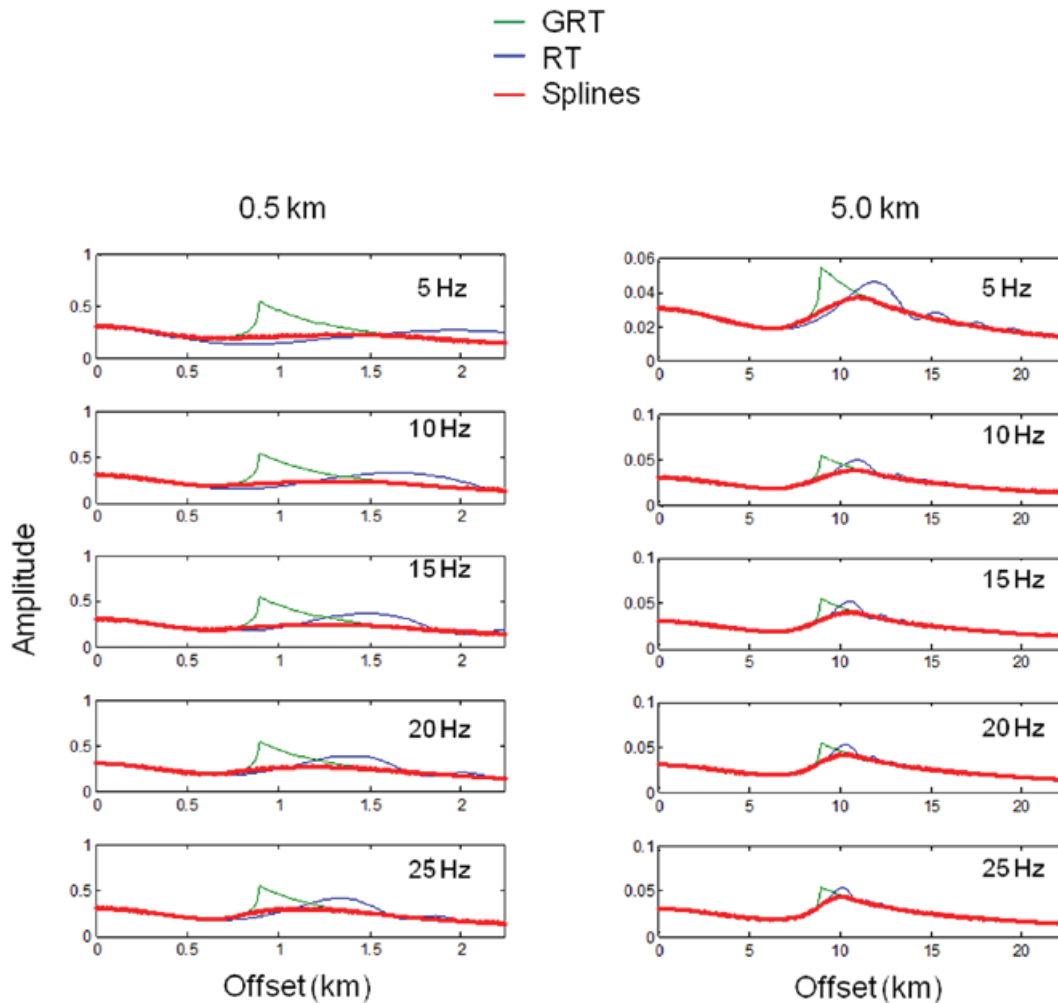
zone. After comparing the AVO curves produced with the exact analytical formulation with ray theory using two half-spaces in contact using numerous geological models, it was observed that smoothing the ray theory amplitudes within the critical zone suffices to match the exact amplitudes very well, independently of the physics of the problem. The fit was obtained by using cubic splines about the critical zone. It was observed that the width of this window decreases exponentially with frequency and with the source-receiver height above the interface. Thus we developed a procedure to determine the window size.



**Figure 3.** AVO curves comparing the exact analytical amplitudes (GRT-blue), geometrical ray theory amplitudes (RT-green) and corrected ray theory amplitudes (CRT-ray) for different frequencies (5-15 Hz) and two different source-receiver height above the interface (0.5 and 5 km). The correction involves the amplitude smoothing within the critical zone  $xz$ . The physical parameters of the model can be found in Table 3.

**Table 3.** Physical parameters of the geological model (see Figure 3).

	$\rho(\text{kg/m}^3)$	$V_p(\text{km/s})$	$V_s(\text{km/s})$
Upper half-space	2.0	2.2	1.3
Lower half-space	2.5	4.0	2.6



**Figure 4.** AVO curves comparing the exact analytical amplitudes (GRT-blue), geometrical ray theory amplitudes (RT-green) and corrected ray theory amplitudes (CRT-ray) for different frequencies (5-15 Hz) and two different source-receiver height above the interface (0.5 and 5 km). The correction involves the amplitude smoothing within the critical zone  $xz$ . The physical parameters of the model can be found in Table 4.

**Table 4.** Physical parameters of the geological model (see Figure 4).

	$\rho(\text{kg/m}^3)$	$V_p(\text{km/s})$	$V_s(\text{km/s})$
Upper half-space	2.0	2.2	1.3
Lower half-space	2.5	3.0	1.5

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